

Optical clocks and geodesy

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Outline

- 1 What is chronometric geodesy?
- 2 Proof-of-principle of chronometric geodesy
- 3 Some definitions and conventions
- 4 ACES/Pharao
- 5 Unifying the gravitational redshift correction
- 6 Chronometric geodesy for high resolution geopotential

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- Chronometry is the science of the measurement of time
- Chronometric geodesy is sometimes named clock-based geodesy

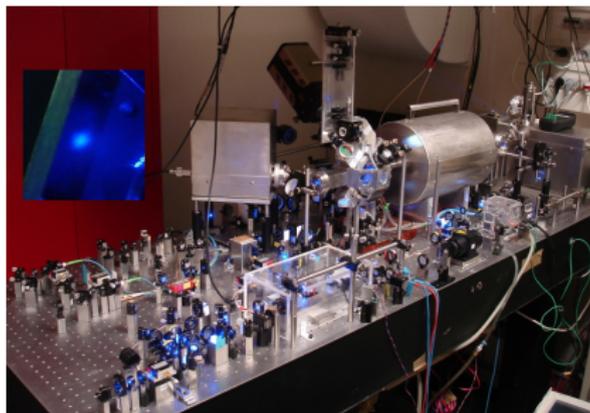


Figure : Strontium clock in SYRTE/Paris Observatory

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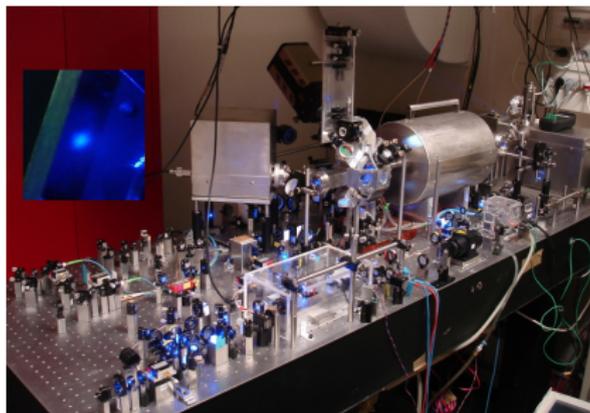


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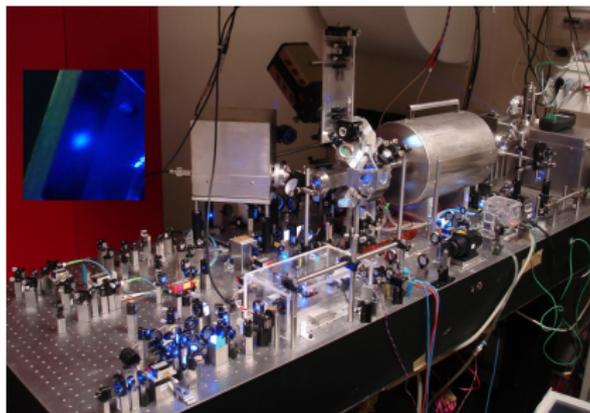


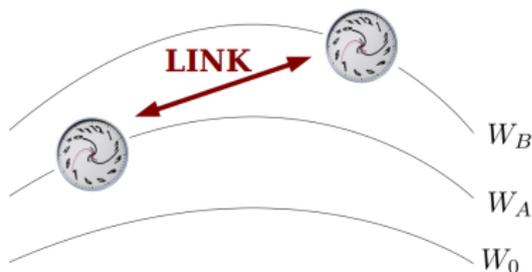
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Basic principle of chronometric geodesy

The flow of time, or the rate of a clock when compared to coordinate time, depends on the **velocity** of the clock and on the **space-time metric** (which depends on the mass/energy distribution).

In the weak-field approximation:

$$\begin{aligned} \frac{\Delta\tau}{\tau} &= \frac{\Delta f}{f} = \frac{U_B - U_A}{c^2} + \frac{v_B^2 - v_A^2}{2c^2} + O(c^{-4}) \\ &= \frac{W_B - W_A}{c^2} + O(c^{-4}) \end{aligned} \quad (1)$$

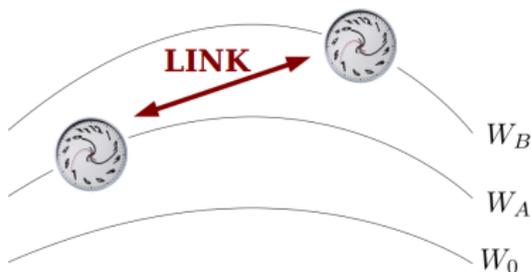


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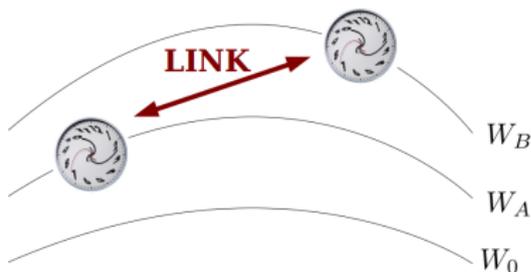
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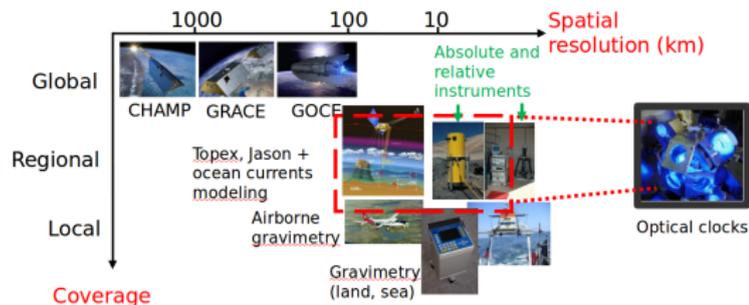
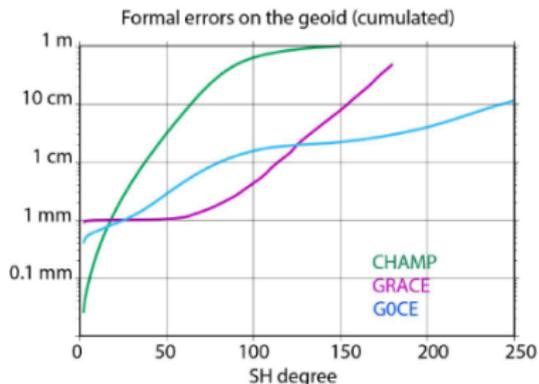
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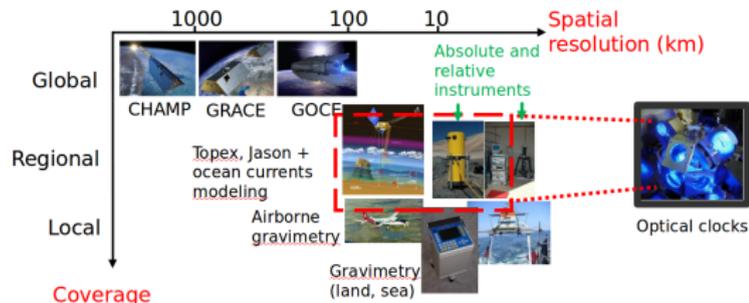
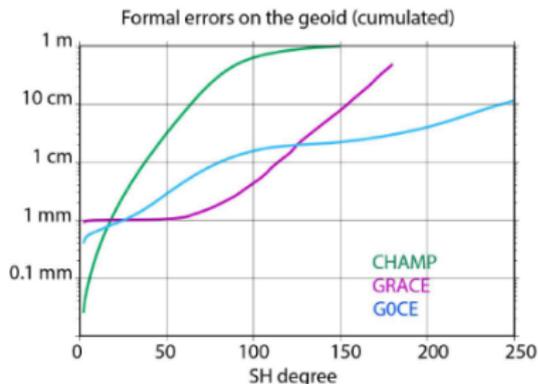
Chronometric observables in geodesy

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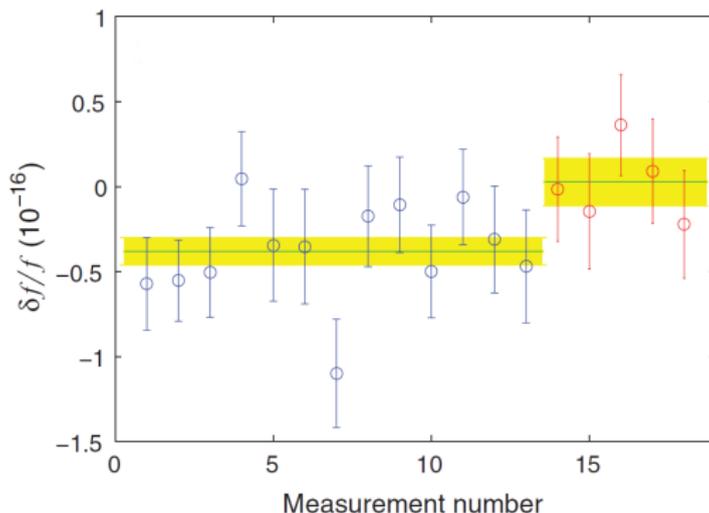
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A local comparison

Experimental demonstration of the dependency of clock frequency with local height [Chou et al., 2010] with two Al^+ optical clocks.

Starting at data point 14, one of the clock is elevated by 33 cm. The net relative shift is measured to be $(41 \pm 16) \times 10^{-18}$.



The shape of the Earth

As a proof-of-principle, one can determine (roughly) J_2 with two clocks:

$$\frac{\Delta f}{f} = \frac{W_B - W_A}{c^2} + O(c^{-4}), \quad W = U + \frac{v^2}{2}$$

$$U = \frac{GM_E}{r} \left[1 + \frac{J_2 R_E^2}{2r^2} (1 - 3 \sin^2(\phi)) \right]$$

- using INRIM CsF1 vs. SYRTE FO2 comparison we find:

$$J_2 = (1.097 \pm 0.016) \times 10^{-3}$$



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Possibilities for technical realisation of a system for **measuring potential differences** over intercontinental distances using clock comparisons
[Vermeer, 1983]

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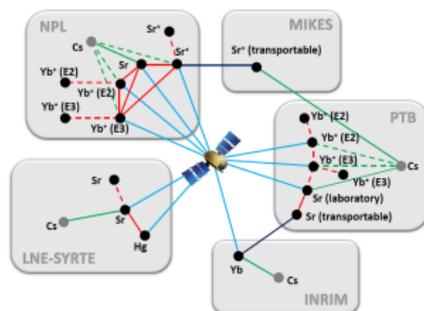
EMRP

European Metrology Research Programme
■ Programme of EURAMET

The EMRP is jointly funded by the EMRP participating countries within EURAMET and the European Union



ITOC International Timescales with Optical Clocks



Main challenge: stable links for frequency comparison

- **Satellite** (GNSS, TWSTFT): intercontinental but limited to $\sim 10^{-16}$, rather long integration time
- **Broadband TWSTFT** (ITOC), **T2L2** (optical): better stability and faster integration, but still far from what is needed

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The chronometric geoid

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Isochronometric surfaces

- An **isochronometric surface** S is a surface where all clocks beat at the same rate:

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- By defining Terrestrial Time (TT) with reference to TCG, the IAU implicitly defined a **reference isochronometric surface** S_0 :

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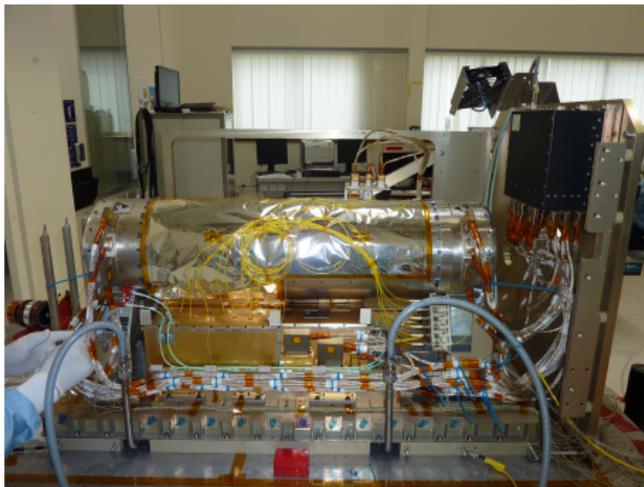
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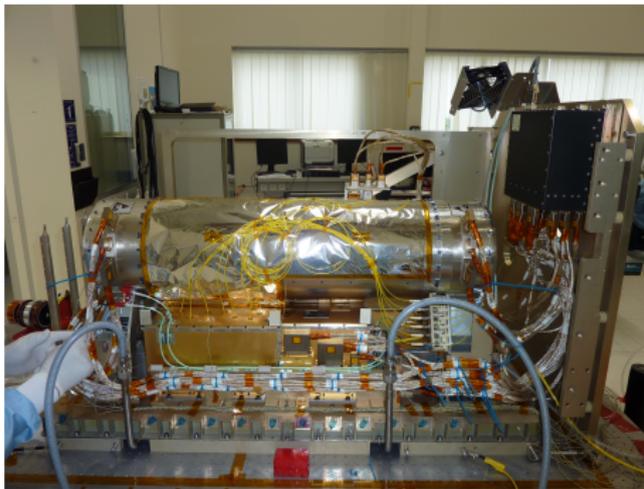
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ACES-Pharao: une fontaine atomique dans l'espace



- Objectifs techniques: créer une échelle de temps de haute exactitude dans l'espace et relier les différentes horloges au sol entre elles
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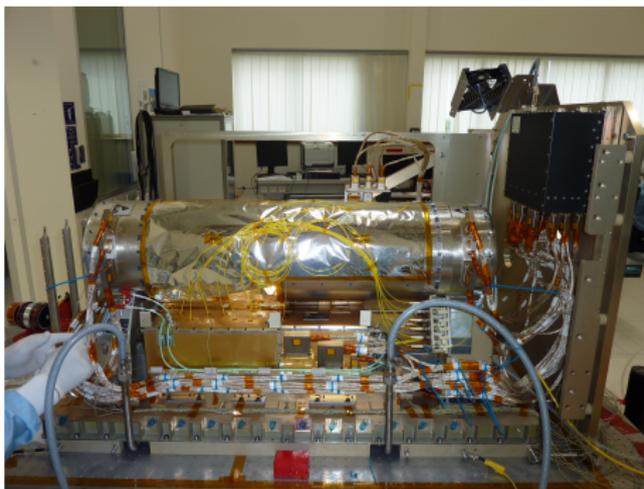
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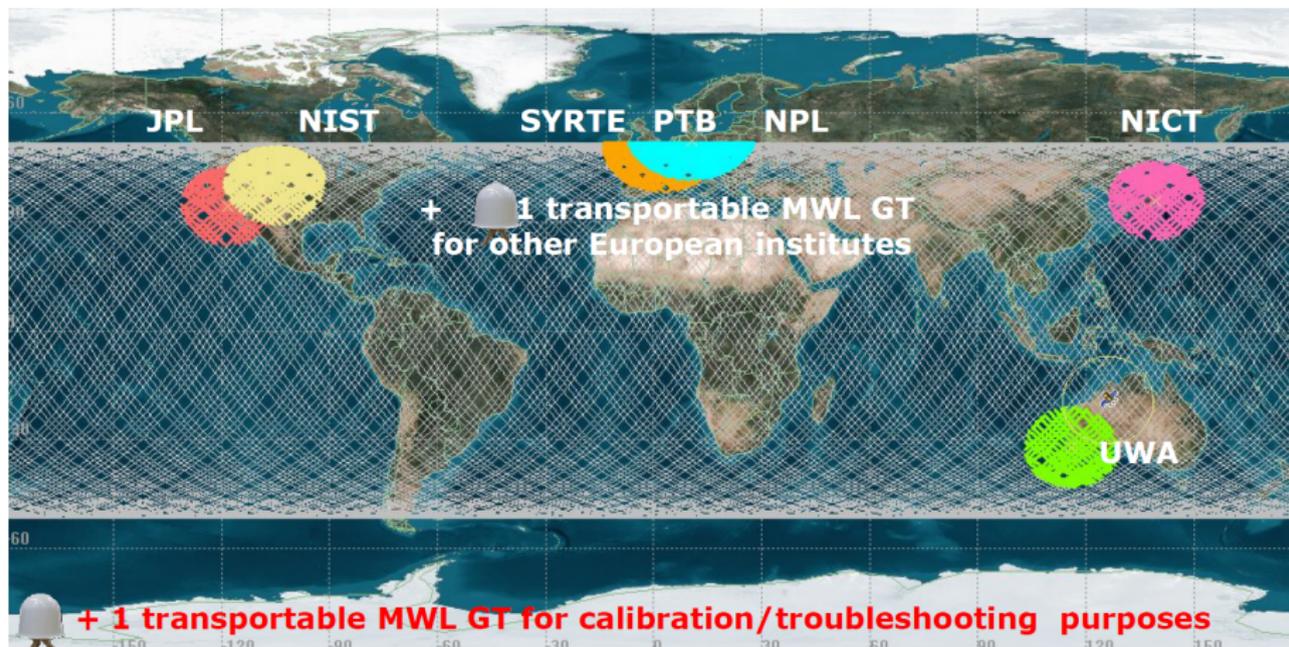


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Les stations au sol



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Consortium of the ITOC EMRP project



National Physical Laboratory (NPL, UK)



Cesky Metrologicky Institut (CMI, Czech Republic)



Istituto Nazionale di Ricerca Metrologica (INRIM, Italy)



Mittatekniikan Keskus (MIKES, Finland)



Physikalisch-Technische Bundesanstalt (PTB, Germany)



SYRTE – Paris Observatory (France)



IfE – Leibniz Universität Hannover (Germany)

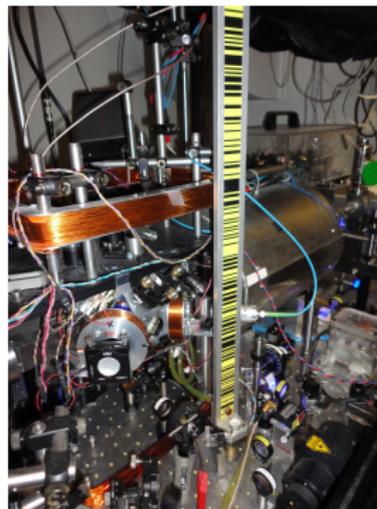
Classical levelling of the clocks [Denker, 2013]

- Design of setups to determine the static gravity potential at all clock locations
- Development of a refined European geoid model including new gravity observations around all relevant clock sites (done by IfE/LUH)



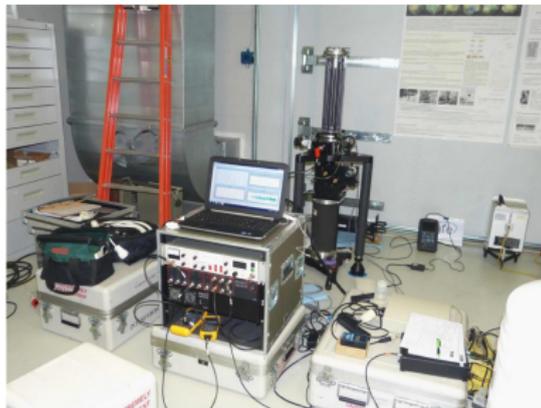
SYRTE clocks
leveling
campaign

(IGN SGN
Travaux
Spéciaux)



Gravity campaigns [Denker, 2015]

NMI	Absolute gravity pts	Relative gravity pts	Comparison with existing pts	Mean difference (mGal)
INRIM	1	35	11	-0.87
LSM	1	122	Italy 10 France 6	+0.20 +0.37
NPL	2	64	25	-0.07
SYRTE	1+2	97	27	-0.12
PTB	1	45	16	+0.08



A new geoid of reference: EGG2015 [Denker, 2015]

NMI	Mean	Std	Min	Max
INRIM	2.7	1.3	0.7	7.3
LSM	3.9	3.9	-5.3	16.6
NPL	-0.9	0.2	-1.4	0.4
OBSPM	-0.7	0.2	-1.4	-0.2
PTB	-0.3	0.1	-0.6	0.0

Table : Difference statistics between the new (EGG2015) and old (EGG2008) quasigeoid heights in centimeters.

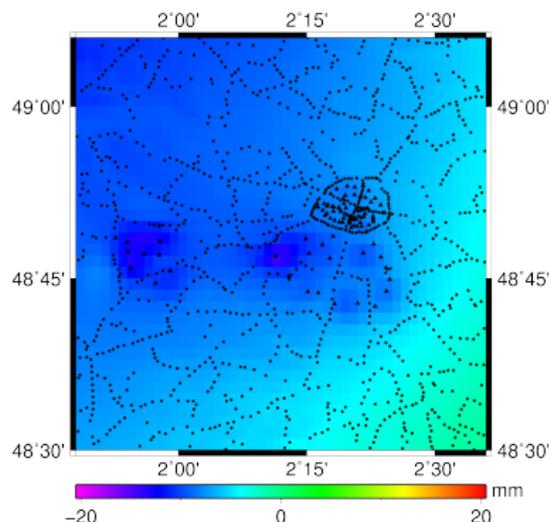


Figure : Differences around OBSPM (\blacktriangle : new ITOC points; \bullet : old points)

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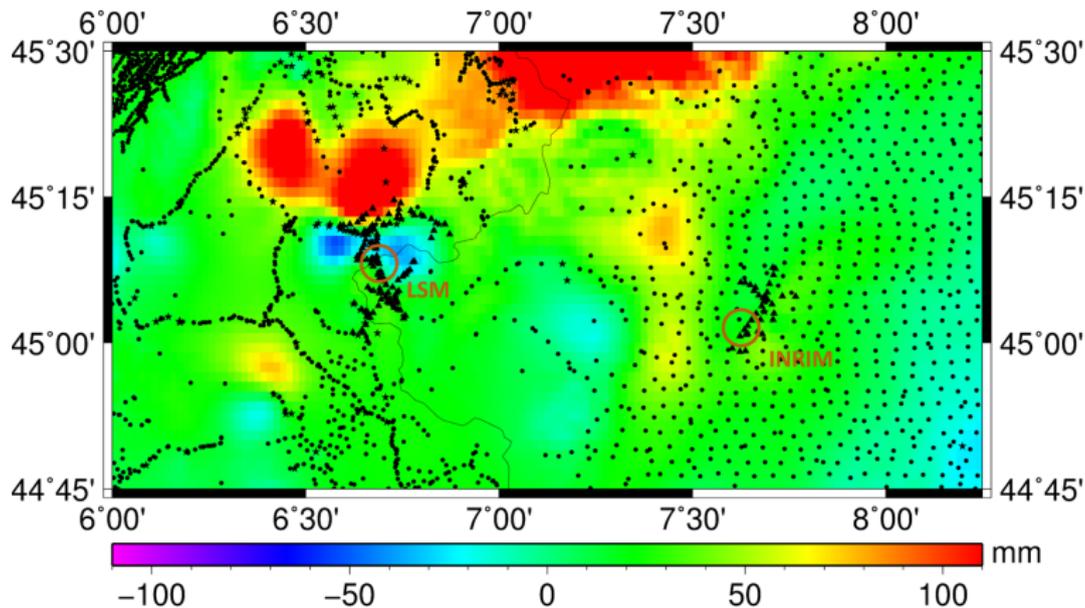
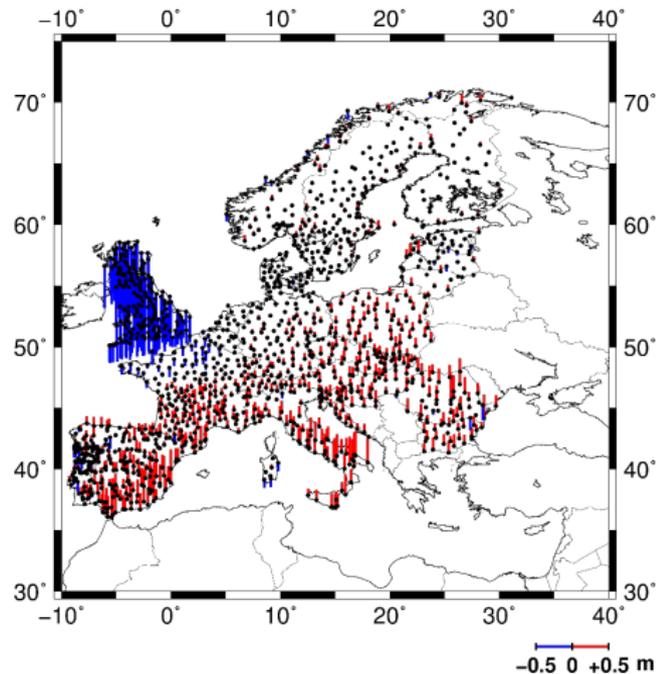


Figure : Differences between the new (EGG2015) and old (EGG2008) quasigeoid heights around INRIM and LSM (\blacktriangle : new ITOC points; \bullet : old points)

Differences between GNSS/geoid & geometric levelling approach [Denker, 2015]

$$\Delta C = C^{(\text{GNSS/geoid})} - C^{(\text{lev})}$$

NMI	ΔC ($10^{-2} \text{ m}^2 \cdot \text{s}^{-2}$)
INRIM	2.3
LSM	-8.3
NPL	-15.3
OBSPM	-11.3
PTB	-2.3



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 - evaluating the contribution of **optical clocks** for the determination of the **geopotential at high spatial resolution**
 - Find the **best locations** to put optical clocks to improve the determination of the geopotential

Chronometric geodesy for high resolution geopotential



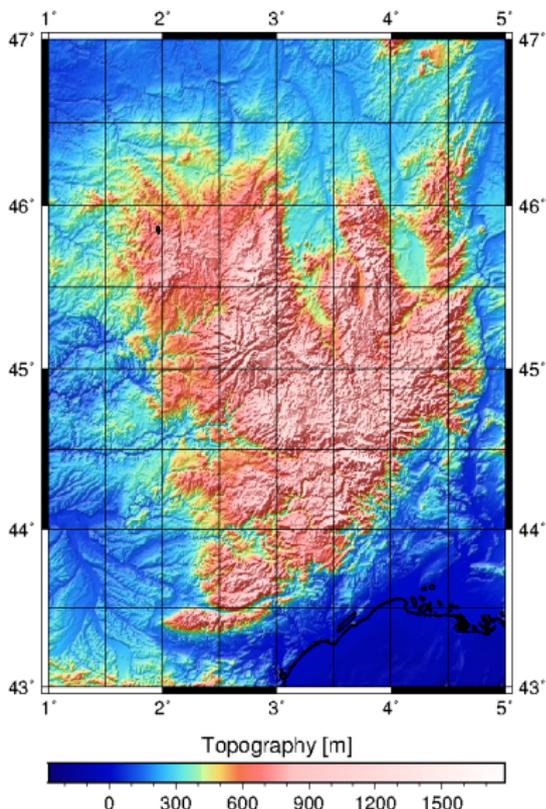
- Collaboration between SYRTE/Obs.Paris, LAREG/IGN and LKB, with the support of GRAM, First-TF and ERC grants
- Goals
 - evaluating the contribution of **optical clocks** for the determination of the **geopotential at high spatial resolution**
 - Find the **best locations** to put optical clocks to improve the determination of the geopotential
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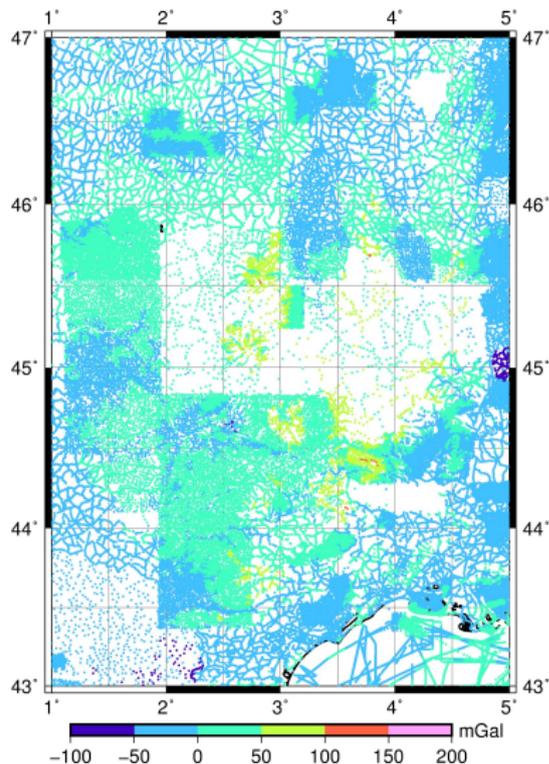
The Auvergne region in France



Interesting region because:

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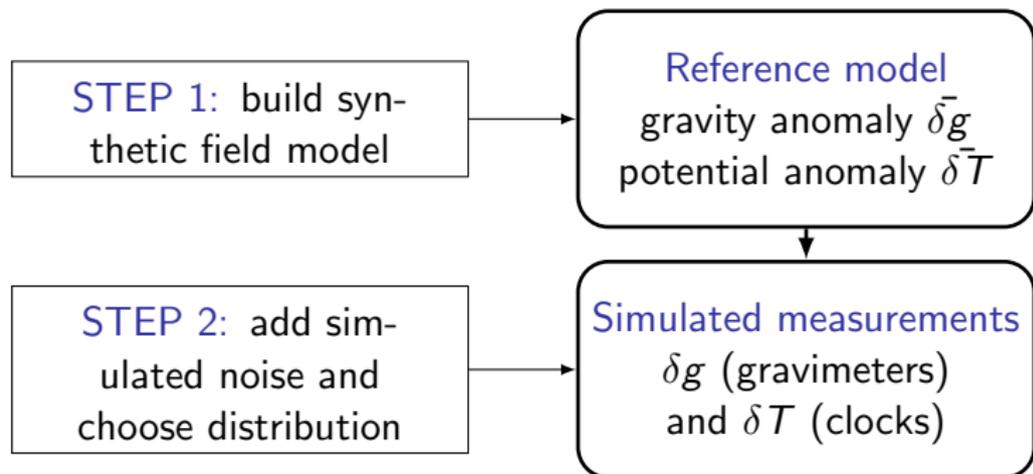
Global methodology

STEP 1: build synthetic field model

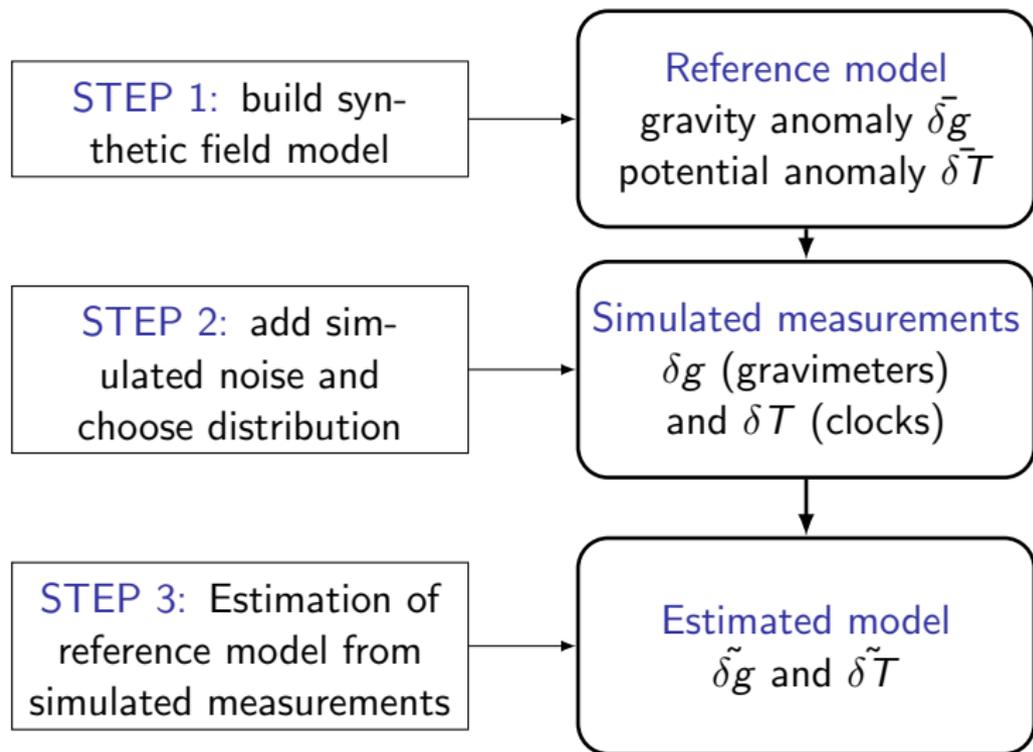
```
graph LR; A[STEP 1: build synthetic field model] --> B[Reference model  
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potential anomaly  $\delta\bar{T}$ ]
```

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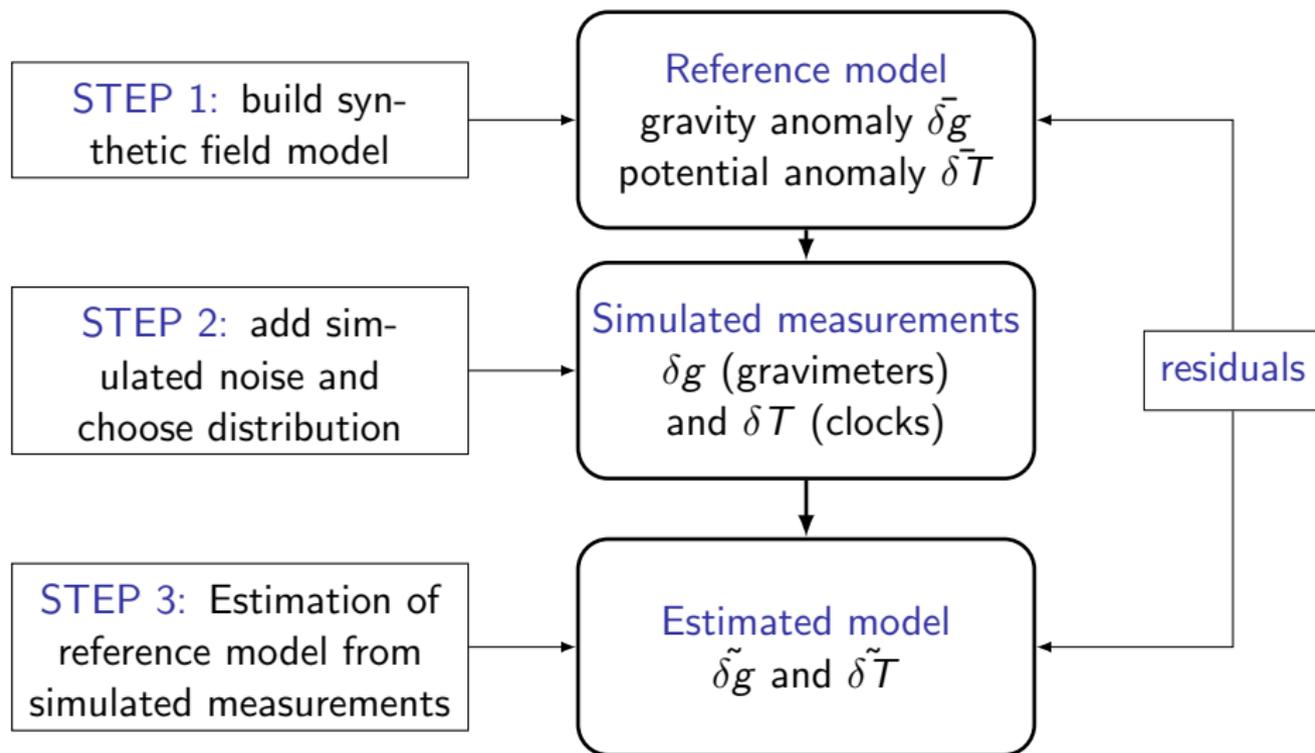
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- 1 Global gravity model at 10 km resolution (EIGEN-6C4, Förste et al. 2014)
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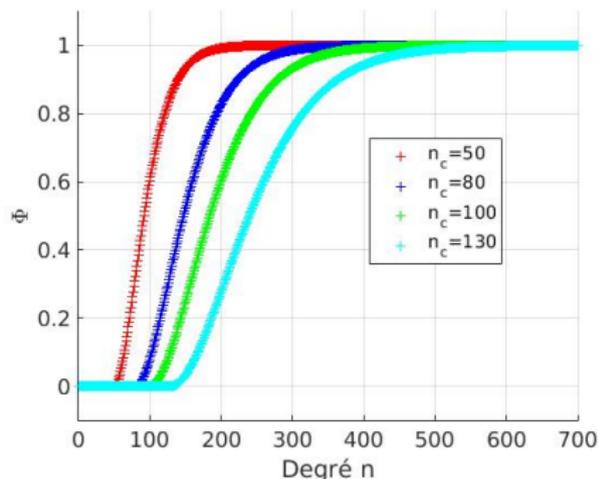


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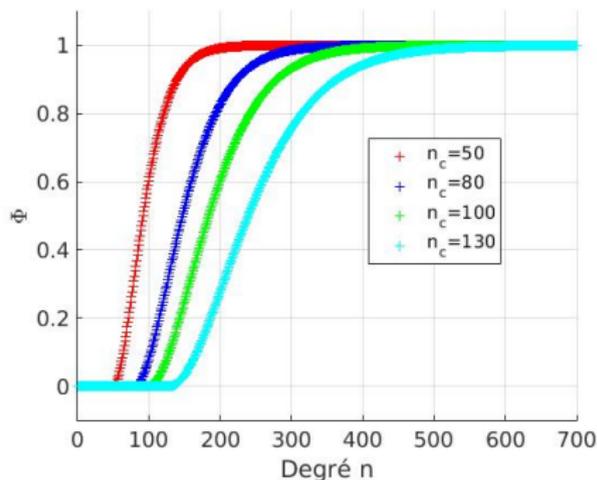


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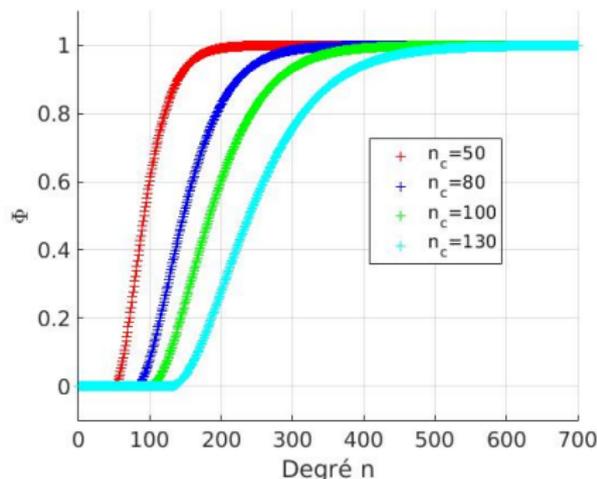


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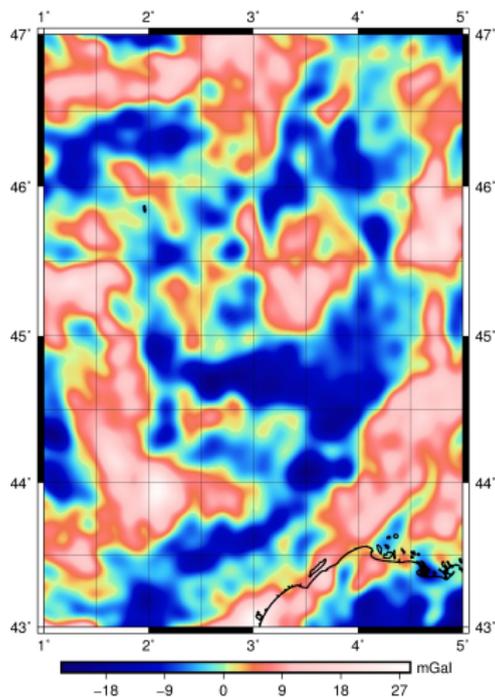


Figure : Reference gravity anomaly

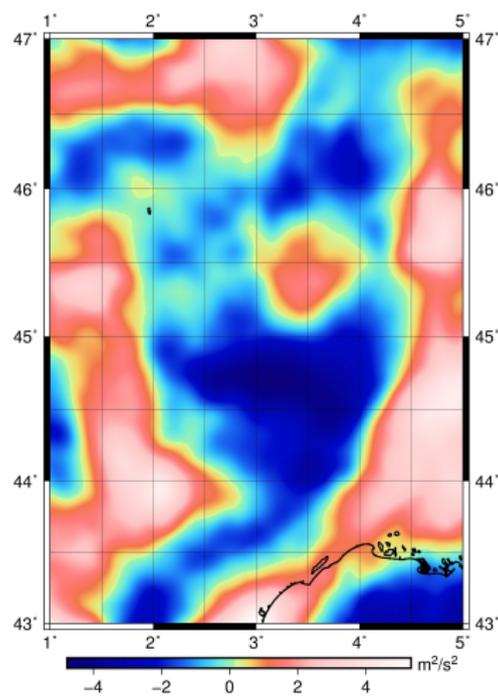
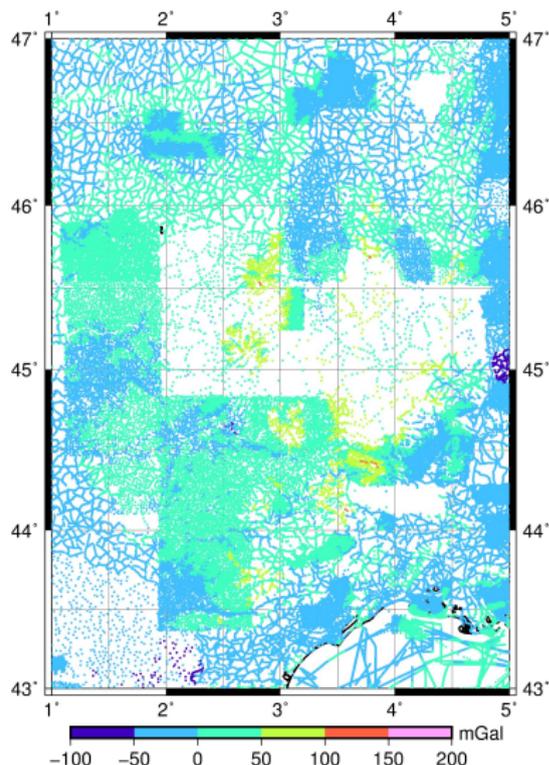


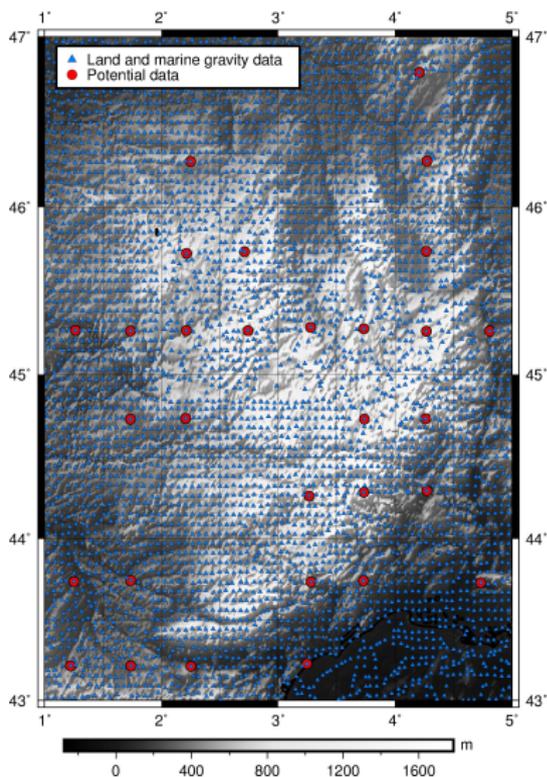
Figure : Reference potential anomaly

STEP 2: add noise and choose observables distribution



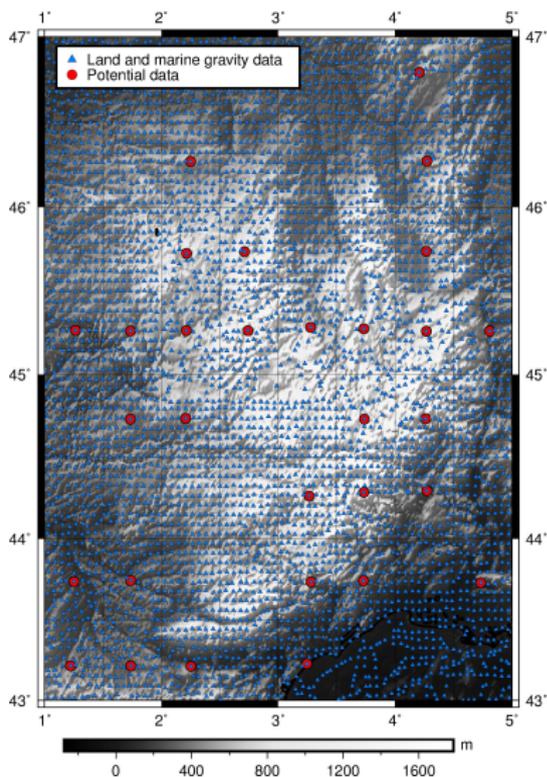
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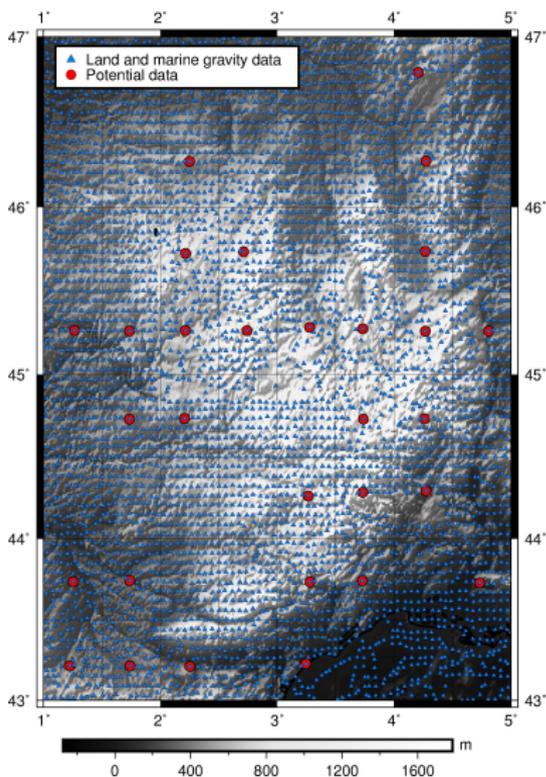
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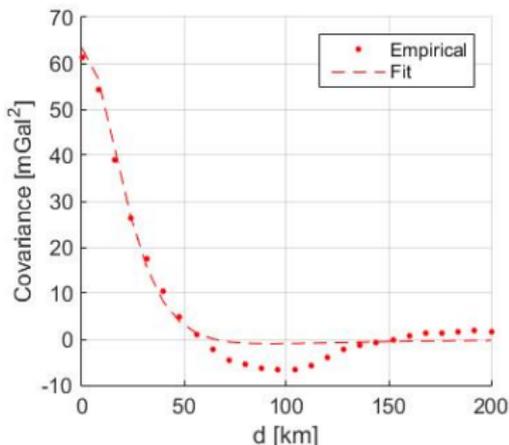


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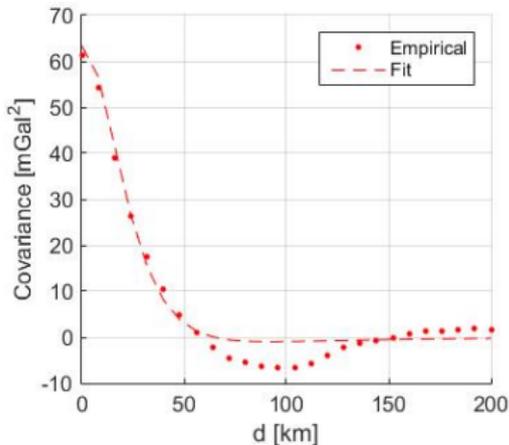


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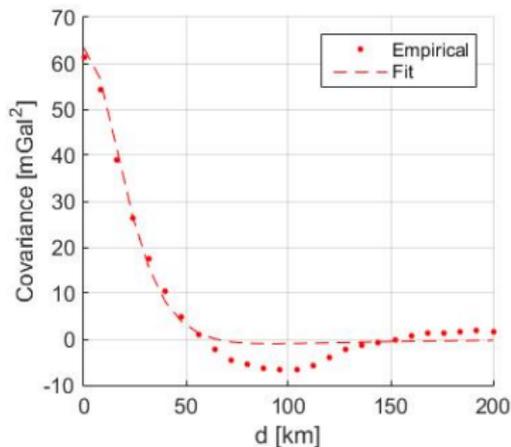


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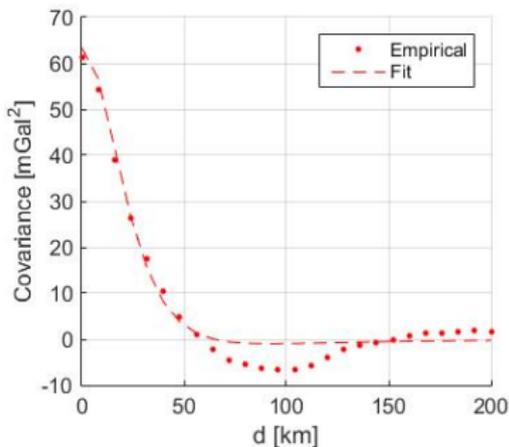
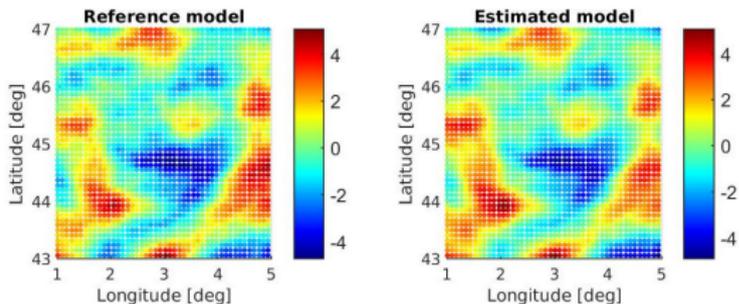


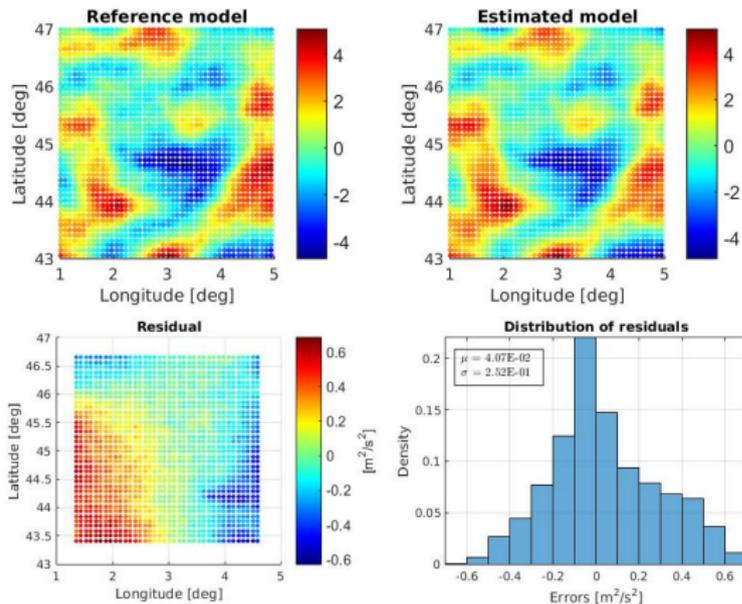
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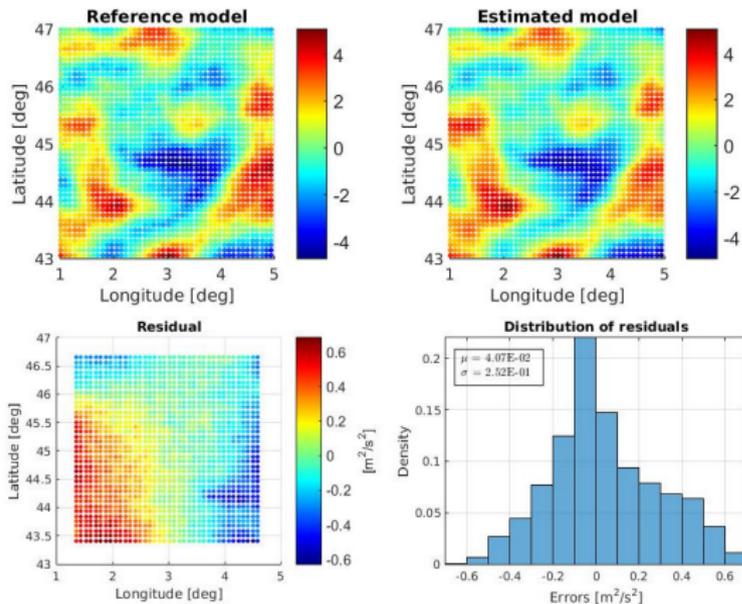
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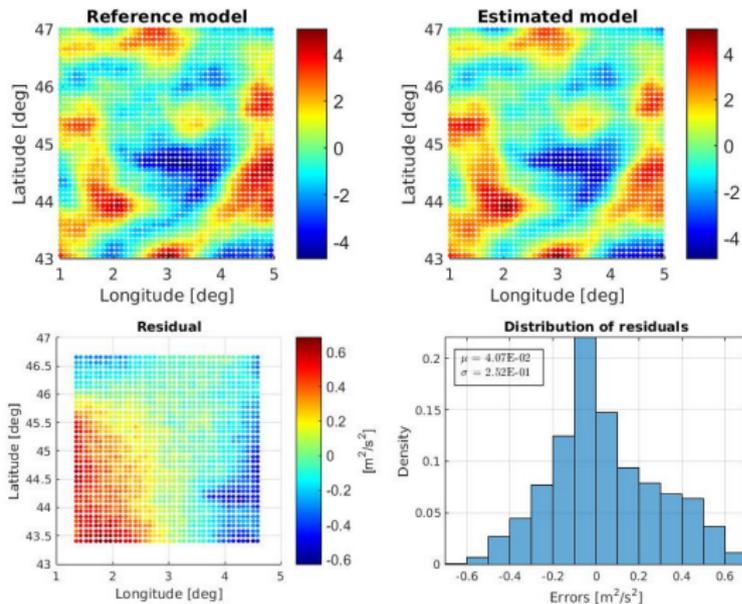
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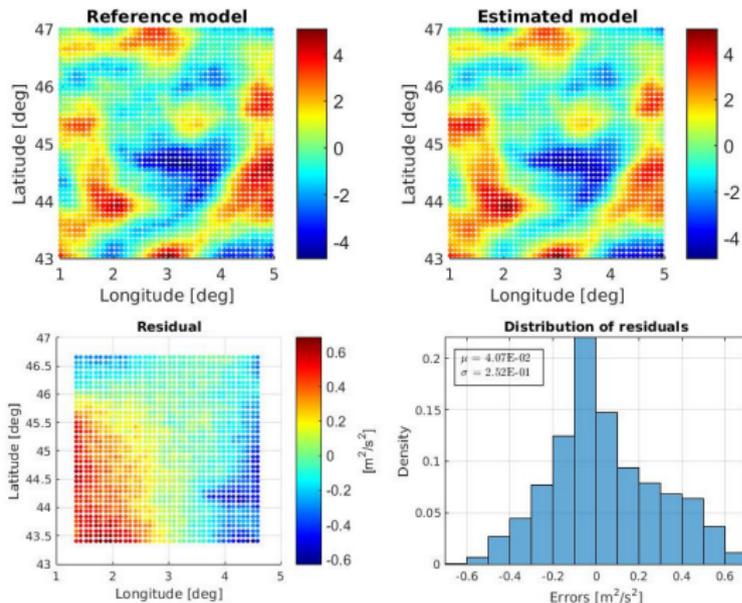
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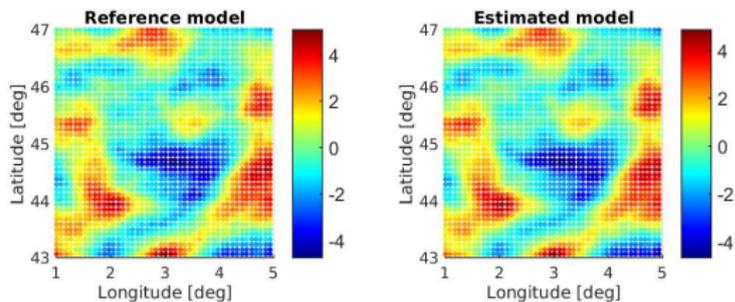
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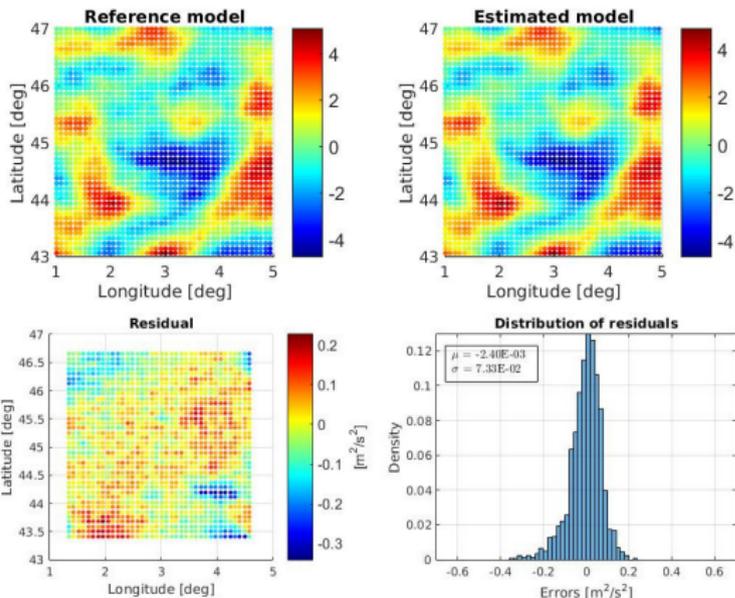
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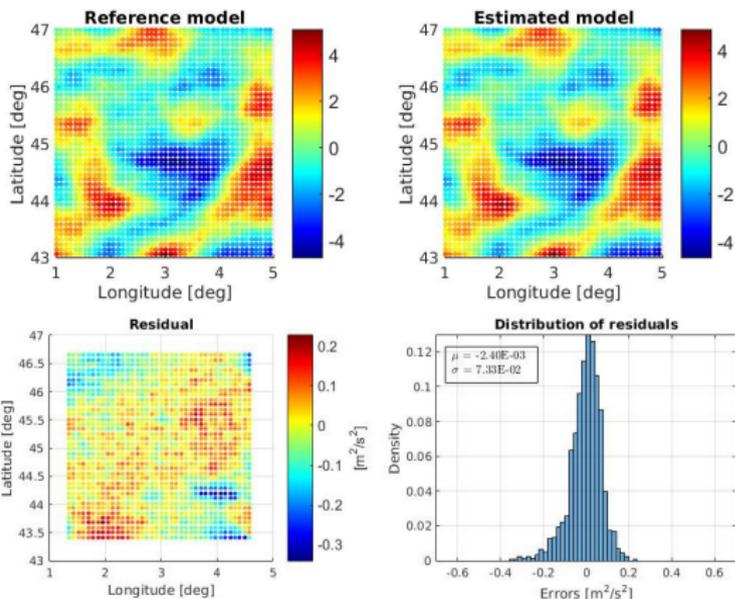
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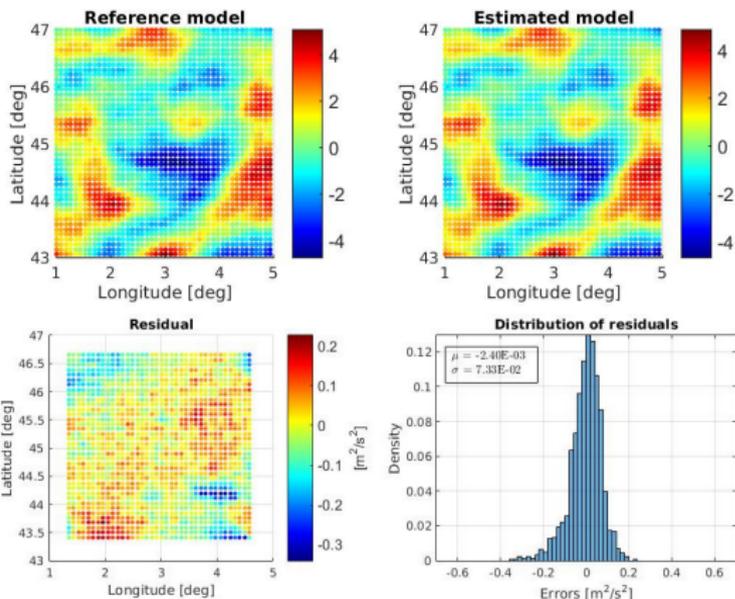
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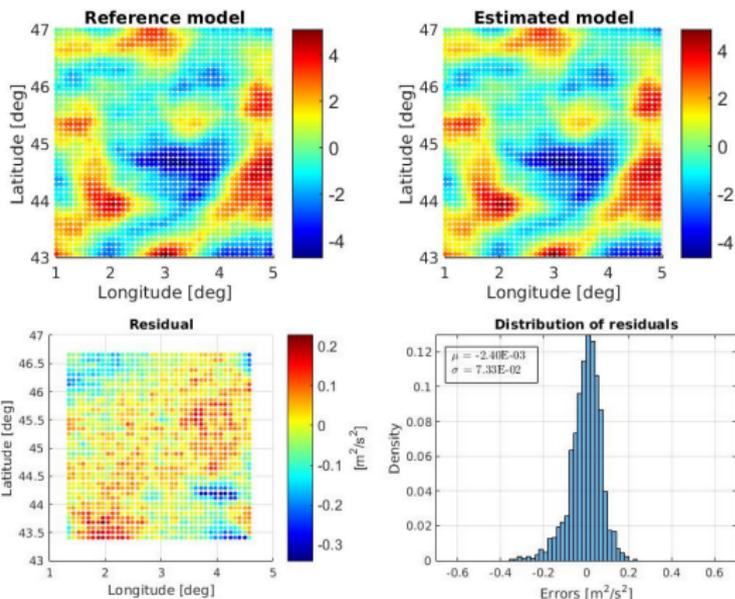
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